



# Radiation effects on mixed convection heat transfer in a vertical square duct

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## Abstract

A numerical study has been made of the interaction of the thermal radiation with laminar mixed convection for a gray fluid (a gas that may have particulates in suspension) in a vertical square duct. Using the vorticity-velocity method, the three-dimensional Navier–Stokes equations and energy equations were solved simultaneously. The integro-differential radiative transfer equation was solved by the discrete ordinates method. Results are presented for a wide range of governing parameters. The effects are emphasized of thermal buoyancy and radiative transfer on the development of velocity and temperature fields, the friction factor and the Nusselt number. The results show that radiation significantly affects the total Nusselt number  $Nu_t$ , and tends to reduce the buoyancy effects. In addition, radiation speeds the development of the temperature field. © 2001 Elsevier Science Ltd. All rights reserved.

## 1. Introduction

Mixed convection in a vertical duct has been of special interest recently due to applications such as solar collectors, heat exchangers, nuclear reactors and electronic equipment. When convection and radiation effects are of similar importance, separate calculation of these and superposition without considering their interaction can result in significant errors. The momentum, energy and radiation transport equations in such cases should be solved simultaneously in order to accurately determine the velocity and temperature fields and heat transfer rates. The present work addresses this problem.

Much work, both theoretical and experimental, has been done on mixed-convection heat transfer in internal flows, as is evident in the review by Jackson et al. [1]. Publications relevant to the present work are briefly reviewed here. Mixed convection heat transfer between vertical parallel plates has been studied by many researchers [2–7]. Aung and Worku [2] show

that the hydrodynamic entry length first increases rapidly with  $Gr/Re$  and then approaches an asymptotic value at large  $Gr/Re$ . Habchi and Acharya [5] find the local Nusselt number to increase with increasing value  $Gr/Re^2$ . Ingham et al. [7] present a numerical method to treat the flow reversal in buoyancy-assisting and -opposing flows. They predict poor heat transfer when flow is retarded by opposed buoyancy, but at large and negative  $Gr/Re$  heat transfer is quite effective.

In flow with low Peclet number or high buoyancy effect, flow reversal may occur. Recently, this case was studied for flows between vertical parallel plates [8–11] or in vertical tubes [12–14]. It appears that at low Peclet number, the omission of axial diffusion terms in the momentum and energy equations leads to serious errors. However, at  $Pe > 50$ , the matching technique used with the boundary-layer equations produces accurate predictions of heat transfer along the heated wall.

Analytical solutions of fully developed mixed convection in vertical rectangular ducts with different heating conditions are presented by Iqbal and Aggarwala [15] and Cheng and Weng [16]. They examined in detail the physical conditions for flow. In practice, the flow typically remains in development for almost the

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Nomenclature			
$a$	width of a square duct, m	$w_o$	inlet mean velocity, m/s
$f$	friction factor, $2\tau_w/(\rho w_o^2)$	$x, y, z$	rectangular coordinate, m
$G, G^*$	dimensional and dimensionless incident radiation, $G^* = G/(4\bar{n}^2\sigma T_w^4)$	$X, Y, Z$	dimensionless rectangular coordinate, $X = x/a, Y = y/a, Z = z/(a \cdot Re)$
$Gr$	Grashof number, $g\beta(T_w - T_o)a^3/\nu^2$	$Z^*$	dimensionless coordinate in the $z$ direction, $Z^* = Z/Pr$
$\bar{h}$	circumferentially averaged heat transfer coefficient, $W/m^2K$	$\nabla^*$	dimensionless gradient operator
$k$	thermal conductivity, $W/m\ k$	<i>Greek symbols</i>	
$m, m'$	direction of the discrete ordinates	$\alpha$	thermal diffusivity, $m^2/s$
$M, N$	number of finite difference divisions in the $X$ and $Y$ directions, respectively	$\beta$	coefficient of thermal expansion
$n$	direction coordinate normal to the duct wall	$\kappa$	extinction coefficient, $m^{-1}$
$\bar{n}$	refractive index	$\kappa_S$	scattering coefficient, $m^{-1}$
$N^*$	order of the phase function	$\epsilon_w$	wall emissivity
$N_c$	conduction-to-radiation parameter, $k\kappa/(4\bar{n}^2\sigma T_w^3)$	$\theta$	dimensionless temperature, $T/T_w$
$Nu_c$	convective Nusselt number	$\theta_o$	dimensionless temperature ratio, $T_o/T_w$
$Nu_r$	radiative Nusselt number	$\mu, \eta, \zeta$	direction cosines
$Nu_t$	total Nusselt number, $Nu_t = Nu_c + Nu_r$	$\nu$	kinematic viscosity, $m^2/s$
$p, P$	dimensional and dimensionless pressure, respectively	$\xi$	dimensionless vorticity in axial direction, $\partial U/\partial Y - \partial V/\partial X$
$\bar{p}$	cross-sectional mean pressure, Pa	$\rho$	density, $kg/m^3$
$\bar{P}$	dimensionless cross-sectional mean pressure	$\sigma$	Stefan–Boltzman constant, $5.67 \times 10^{-8} W/m^2 K^{-4}$
$P'$	perturbation term about the mean pressure	$\tau$	optical thickness
$\bar{P}$		$\tau_w$	wall shear stress, kPa
$Pe$	Peclet number, $Pr \cdot Re$	$\psi$	dimensionless radiation intensity, $\pi I/(\bar{n}^2\sigma T_w^4)$
$P_n$	Legendre polynomial	$\phi$	scattering phase function
$Pr$	Prandtl number, $\nu/\alpha$	$\omega$	single scattering albedo
$\vec{q}_c$	convective heat flux	$\vec{\Omega}, \vec{\Omega}'$	outward and inward direction of radiation
$\vec{q}_r$	radiation heat flux	$\Omega$	solid angle
$\vec{q}_t$	total heat flux	<i>Subscripts</i>	
$\vec{Q}_r$	dimensionless radiation flux	b	bulk fluid quantity
$Ra$	Rayleigh number $Pr \cdot Gr$	c	convective
$Re$	Reynolds number, $w_o a/\nu$	o	condition at inlet
$T$	temperature, K	r	radiative
$u, v, w$	velocity components in $x, y$ and $z$ directions, respectively, m/s	w	value at wall
$U, V, W$	dimensionless velocity components in $X, Y$ and $Z$ directions, respectively	<i>Superscript</i>	
		-	averaged value

full length of the duct, so a knowledge of the developing mixed convection is important. Cheng and Weng [17] present relevant numerical results of developing mixed convection heat transfer in vertical rectangular ducts without consideration of the thermal radiation effects.

Forced convection in a channel with radiation effect only (i.e., without thermal buoyancy) has been examined by several investigators [18–24] during the past thirty years for ducts with a prescribed heat flux or wall temperature distribution on the wall surfaces.

In these studies, various approximate methods were used for radiative transport in the medium (the flowing fluid). Among these, the differential approximation methods describe approximation variation of the intensity of radiation as function of position and angle.

No numerical studies have been published of combined convection and radiation in mixed convection duct flows. This motivates the present study. In addition, flow is considered to be hydrodynamically and thermally developing.

## 2. Analysis

As a preliminary attempt to study thermal radiation effects, flow is assumed to be laminar. Practical flows are, of course, often turbulent. Accounting for turbulence, however, greatly complicates the simulation. The present work is thus a first step.

Consider, therefore, steady laminar flow through a vertical square duct of width  $a$  and uniform wall Temperature  $T_w$ . A uniform inlet axial velocity  $w_o$  and a uniform inlet temperature  $T_o$  are imposed at the entrance  $z = 0$ . The  $u$ ,  $v$ , and  $w$  are the velocity components in the  $x$ ,  $y$ , and  $z$  directions, respectively. To simplify analysis, the fluid properties are taken to be constants expect for the density variation in the buoyancy term of the momentum equation. The Boussinesq approximation is invoked for the thermal buoyancy effect. The viscous dissipation and compression effects in the energy equation are negligible due to low Mach number flows. Additionally, the gas is assumed to be gray, absorbing, emitting, and scattering.

The flow is treated as parabolic and a space-averaged pressure  $\bar{p}$  is imposed in the momentum equations to prevail at each cross-section, thus permitting a decoupling from the pressure  $p_m$  in the cross-stream momentum equations. This “pressure uncoupling” follows the parabolic-flow practice and, together with the assumption that neither momentum nor heat is diffused in the axial direction (supported by an order of magnitude analysis), permits a marching-integration calculation procedure [25]. To conveniently present the governing equations, the pressure  $p$  can be represented as the sum of a cross-section mean pressure  $\bar{p}(z)$ , which drives the main flow, and a perturbation about the mean,  $p'(x, y, z)$ , which drives the cross-stream flow,

$$p = \bar{p}(z) + p'(x, y, z). \quad (1)$$

By introducing a vorticity function in the axial direction,  $\xi = \partial U/\partial Y - \partial V/\partial X$ , the vorticity–velocity formulation of Navier–Stokes equations is obtained:

$$\partial^2 U/\partial X^2 + \partial^2 U/\partial Y^2 = \partial \xi/\partial Y - \partial^2 W/\partial X \partial Z, \quad (2)$$

$$\partial^2 V/\partial X^2 + \partial^2 V/\partial Y^2 = -\partial \xi/\partial X - \partial^2 W/\partial Y \partial Z, \quad (3)$$

$$\begin{aligned} U \partial \xi/\partial X + V \partial \xi/\partial Y + W \partial \xi/\partial Z + \xi(\partial U/\partial X + \partial V/\partial Y) \\ + (\partial W/\partial Y \cdot \partial U/\partial Z - \partial W/\partial X \cdot \partial V/\partial Z) \\ = \partial^2 \xi/\partial X^2 + \partial^2 \xi/\partial Y^2, \end{aligned} \quad (4)$$

$$\begin{aligned} U \partial W/\partial X + V \partial W/\partial Y + W \partial W/\partial Z \\ = -d\bar{P}/dZ + Ra/(Pr \cdot Re) \cdot (\theta - \theta_o)/(1 - \theta_o) \\ + \partial^2 W/\partial X^2 + \partial^2 W/\partial Y^2, \end{aligned} \quad (5)$$

$$\begin{aligned} U \partial \theta/\partial X + V \partial \theta/\partial Y + W \partial \theta/\partial Z \\ = [\partial^2 \theta/\partial X^2 + \partial^2 \theta/\partial Y^2 + (1 - \omega)\tau^2/N_c \\ \cdot (G^* - \theta^4)]/Pr. \end{aligned} \quad (6)$$

In the above formulation, the dimensionless groups are defined as follows:

$$\begin{aligned} X = x/a, \quad Y = y/a, \\ Z = z/(a \cdot Re), \quad U = ua/v, \\ V = va/v, \quad W = w/w_o, \\ \bar{P} = \bar{p}/(\rho w_o^2), \quad \theta = T/T_w, \\ \theta_o = T_o/T_w, \quad Re = w_o a/v, \\ Pr = \nu/\alpha, \quad Gr = g\beta(T_w - T_o)a^3/\nu^2, \\ N_c = k\kappa/(4\bar{n}^2\sigma T_w^3), \quad \tau = \kappa a, \\ \omega = \kappa_s/\kappa, \quad G^* = G/(\bar{n}^2\sigma T_w^4), \\ Ra = Pr \cdot Gr, \quad Z^* = Z/Pr. \end{aligned} \quad (7)$$

The overall mass flow rate at every axial location must be balanced in the duct flow, i.e.,

$$\int_0^1 \int_0^1 W \, dX \, dY = 1. \quad (8)$$

For the radiation part of this problem, we consider a gray, absorbing and scattering medium. The radiation transfer equation is given in the dimensionless form as

$$\begin{aligned} \mu \partial \psi/\partial X + \eta \partial \psi/\partial Y + \tau \psi \\ = (1 - \omega)\tau\theta^4 + \frac{\omega\tau}{4\pi} \int_{\Omega'=4\pi} \phi(\vec{\Omega}', \vec{\Omega}) \psi \, d\Omega', \end{aligned} \quad (9)$$

where the term  $\partial \psi/\partial Z$  is omitted under the assumption of negligible axial radiation, i.e.,  $\partial q_{rz}/\partial Z \ll \partial q_{rx}/\partial X + \partial q_{ry}/\partial Y$ . This assumption has been corroborated independently by Echigo et al. [20] and Campo and Schuler [23] who showed that the axial radiation penetrates one or two diameters upstream of the origin of the heat exchange section. In a dimensionless form, this penetrated distance is about  $Z^* = 0.001$  for the flow with  $Re = 2000$  in this work. Therefore, the axial radiation is discarded. In Eq. (9),  $\psi = \pi I/(\bar{n}^2\sigma T_w^4)$  is the dimensionless intensity of radiation at a point  $(X, Y)$  in the direction  $\vec{\Omega}$  defined by the direction cosines  $\mu, \eta$ , and  $\zeta$ ,  $\omega$  is the single scattering albedo, and  $\phi(\vec{\Omega}', \vec{\Omega})$  is the scattering phase function, which is expressed in terms of Legendre polynomials as

$$\phi(\vec{\Omega}', \vec{\Omega}) = \sum_{n=0}^{N^*} a_n P_n(\mu'\mu + \eta'\eta + \zeta'\zeta). \quad (10)$$

If the surface is assumed to be opaque, gray and diffusely reflecting, the wall of boundary condition can be written as

$$\begin{aligned} \psi_w(\vec{\Omega}) = \epsilon_w + \frac{(1 - \epsilon_w)}{\pi} \int_{\vec{n} \cdot \vec{\Omega}' < 0} |\vec{n} \cdot \vec{\Omega}'| \\ |\psi_w(\vec{\Omega}')| \, d\Omega', \quad \vec{n} \cdot \vec{\Omega} > 0, \end{aligned} \quad (11)$$

where  $\epsilon_w$  is the wall emissivity, and  $\vec{n}$  is the unit normal vector pointing out of the duct wall into the medium.

Once the dimensionless radiation intensity  $\psi$  is known, the dimensionless radiation flux vector and incident radiation are found from

$$\vec{Q}_r = \vec{q}_r / (4\bar{n}^2 \sigma T_w^4) = (1/4\pi) \cdot \int_{\Omega=4\pi} \vec{\Omega} \psi \, d\Omega, \quad (12)$$

$$G^* = G / (4\bar{n}^2 \sigma T_w^4) = (1/4\pi) \cdot \int_{\Omega=4\pi} \psi \, d\Omega. \quad (13)$$

The boundary conditions for the convective governing equations are

$$U = V = W = 0, \quad \theta = 1 \text{ at the duct walls,} \quad (14)$$

$$W = 1, \quad U = V = \xi = 0, \quad \theta = \theta_o \text{ at the entrance } Z = 0. \quad (15)$$

After the velocity and temperature fields are obtained, computation of the circumferentially averaged friction factor and Nusselt number is of practical interest. The dimensionless friction factor is

$$fRe = 2(\overline{\partial W / \partial n})_w. \quad (16)$$

Energy transport from the duct wall to the gas flow in the presence of thermal radiation depends on two related factors: The fluid temperature gradient on the duct wall and the rate of radiative heat exchange.

$$q_t = q_c + q_r = -k(\partial T / \partial n)_w + q_r. \quad (17)$$

The local total Nusselt number  $Nu_t$  defined as

$$Nu_t = \bar{h}a/k = q_t a / [k(T_w - T_b)], \quad (18)$$

consists of both convection and radiation heat transfer rates. It is written as

$$Nu_t = Nu_c + Nu_r, \quad (19)$$

where  $Nu_c$  and  $Nu_r$  are respectively, the local convective Nusselt number and radiative Nusselt number, and are defined as

$$Nu_c = -(\overline{\partial \theta_w / \partial n}) / (1 - \theta_b) \quad (20)$$

and

$$Nu_r = (\tau \vec{Q}_r / N_c) / (1 - \theta_b), \quad (21)$$

where the overbar means the circumferential average,  $\vec{Q}_r$  is the dimensionless radiative heat flux at the duct wall, and the bulk temperature  $\theta_b$  is

$$\theta_b = \int_0^1 \int_0^1 \theta \cdot W \, dX dY. \quad (22)$$

The parameters involved are the Prandtl number  $Pr$ , the ratio of the Rayleigh to Reynolds numbers,  $Ra/Re$ , the conduction-to-radiation  $N_c$ , the optical thickness  $\tau$ ,

the scattering albedo  $\omega$ , the wall emissivity  $\epsilon_w$  and temperature ratio of inlet fluid and wall  $\theta_o$ . The  $Ra/Re$  measures the importance of the thermal buoyancy effects relative to the inertia force. The conduction-to-radiation parameter  $N_c$  characterizes the relative importance with respect to radiation. The effect of radiation is getting strong as  $N_c$  decreases. To reduce the computational effort, the temperature ratio of inlet fluid and wall is fixed to be  $\theta_o = 0.3$  and the Prandtl number is set to be  $Pr = 0.7$ . Effects of other parameters are examined in detail.

### 3. Solution method

The governing equations are solved numerically by the vorticity–velocity method for three-dimensional parabolic flow [25,26]. For a given combination of parameters, the field solutions are calculated by a marching technique based on the Du Fort–Frankel scheme [27]. Details of the solution procedure have been described elsewhere [25,26] and are not repeated herein. In the present work, the radiative transfer equation is solved by the discrete ordinates method with  $S_N$  quadrature [28–30]. The solid angle  $4\pi$  is discretized over a finite number of directions, the radiative transfer equation is applied at these directions with the integral term replaced by a quadrature. The discrete ordinates representation of the radiative transfer equation is

$$\begin{aligned} \mu_m \partial \psi_m / \partial X + \eta_m \partial \psi_m / \partial Y + \tau \psi_m \\ = (1 - \omega) \tau \theta^4 + \frac{\omega \tau}{4\pi} \sum_{m'} w_{m'}^* \phi_{m'm} \psi_{m'}, \end{aligned} \quad (23)$$

where subscripts  $m$  and  $m'$  represent the directions of the discrete ordinates, and the  $w_m^*$  are the quadrature weights. The associated boundary conditions are

$$\begin{aligned} \psi_m = \epsilon_w + \frac{(1 - \epsilon_w)}{\pi} \sum_{m'} \psi_{m'} |\mu_{m'}| w_{m'}^*, \\ \mu_m > 0, \quad \mu_m < 0, \quad X = 0, \end{aligned} \quad (24)$$

$$\begin{aligned} \psi_m = \epsilon_w + \frac{(1 - \epsilon_w)}{\pi} \sum_{m'} \psi_{m'} |\mu_{m'}| w_{m'}^*, \\ \mu_m < 0, \quad \mu_{m'} > 0, \quad X = 1, \end{aligned} \quad (25)$$

$$\begin{aligned} \psi_m = \epsilon_w + \frac{(1 - \epsilon_w)}{\pi} \sum_{m'} \psi_{m'} |\mu_{m'}| w_{m'}^*, \\ \eta_m > 0, \quad \eta_{m'} < 0, \quad Y = 0, \end{aligned} \quad (26)$$

$$\begin{aligned} \psi_m = \epsilon_w + \frac{(1 - \epsilon_w)}{\pi} \sum_{m'} \psi_{m'} |\mu_{m'}| w_{m'}^*, \\ \eta_m < 0, \quad \eta_{m'} < 0, \quad Y = 1. \end{aligned} \quad (27)$$

The discrete form of the phase function  $\phi_{m'm}$  is

$$\phi_{m'm} = \sum_{n=0}^{N^*} a_n P_n(\mu_{m'} \mu_m + \eta_{m'} \eta_m + \zeta_{m'} \zeta_m). \quad (28)$$

Table 1  
The discrete directions and quadrature weights for the  $S_6$  method (one octant only)

$m$	$\mu_m$	$\eta_m$	$\zeta_m$	$w_m^*$
1	0.948235	0.224556	0.224556	$\pi/6$
2	0.689048	0.689048	0.224556	$\pi/6$
3	0.224556	0.948235	0.224556	$\pi/6$
4	0.689048	0.224556	0.689048	$\pi/6$
5	0.224556	0.689048	0.689048	$\pi/6$
6	0.224556	0.224556	0.948235	$\pi/6$

Table 2  
Comparisons of local  $Nu_t$  for various grid arrangements for  $Ra/Re = 100, N_c = 0.05, \tau = 1, \omega = 0, \epsilon_w = 0.5$  and  $\theta_o = 0.3$

$M \times N$ ( $\Delta Z^*$ )	$Z^*$					
	0.001	0.005	0.01	0.05	0.1	0.3
$51 \times 51$ ( $1 \times 10^{-5} - 5 \times 10^{-4}$ )	15.528	9.533	8.274	8.460	8.787	8.398
$81 \times 81$ ( $1 \times 10^{-5} - 5 \times 10^{-4}$ )	15.232	9.428	8.275	8.456	8.732	8.408
$51 \times 51$ ( $1 \times 10^{-6} - 5 \times 10^{-4}$ )	15.640	9.397	8.275	8.464	8.787	8.463
$31 \times 31$ ( $1 \times 10^{-5} - 5 \times 10^{-4}$ )	17.241	9.484	8.265	8.442	8.790	8.608

The angular discretization transforms the integro-differential radiative transfer equation into a set of coupled partial differential equations. The discrete directions and quadrature weights used in this work are based on the momentum-matching technique [31], as listed in Table 1. The total number of discrete directions is 24 when the  $S_6$  scheme is employed for a two-dimensional geometry.

To ensure the independence of the grid points used in this work, a numerical experiment was made on the grids in the cross-sectional plane ( $M \times N$ ) and axial step size ( $\Delta Z^*$ ). In this study, grids were arranged to be uniform in the cross-sectional direction but non-uniformly distributed in the axial direction for the uneven variations of field properties in the entrance region. It is found, Table 2, that the deviations in local total Nusselt number  $Nu_t$  calculated with  $M \times N = 51 \times 51$  and  $81 \times 81$  ( $\Delta Z^* = 1 \times 10^{-5} - 5 \times 10^{-4}$ ) are always less than 2% for  $Ra/Re = 100, N_c = 0.05, \tau = 1, \omega = 0$  and  $\epsilon_w = 0.5$ . Furthermore, the deviations in  $Nu_t$  calculated using  $M \times N$  ( $\Delta Z^* = 51 \times 51$  ( $\Delta Z^* = 1 \times 10^{-6} - 5 \times 10^{-4}$ ) and  $51 \times 51$  ( $\Delta Z^* = 1 \times 10^{-5} - 5 \times 10^{-4}$ ) are all less than 2%. Accordingly, the computations involving a  $M \times N$  ( $\Delta Z^* = 51 \times 51$  ( $\Delta Z^* = 1 \times 10^{-5} - 5 \times 10^{-4}$ ) grid are considered to be sufficiently accurate to describe the flow and heat transfer in a vertical square duct. All the results presented in the next section are computed using the latter grid. As a partial verification of the computational procedure, results were initially obtained for convection heat transfer in a vertical square duct without thermal

radiation effect and compared with those of Cheng and Weng [17]. The Nusselt number and friction factor were found to agree within 2%. Through these numerical tests, the present numerical scheme was suitable for the present problem.

#### 4. Results and discussion

The developing axial velocity profiles along the centerline  $X = 0.5$  at various conditions are shown in Fig. 1, where curves A, B, C, D, and E denote profiles at  $Z^* = 0.001, 0.01, 0.05, 0.1,$  and  $0.3,$  respectively. For comparison, the results without thermal radiation effect are also included in Fig. 1, as indicated by the dashed lines. It should be mentioned that the uniform velocity at the entrance  $Z = 0$  corresponds to  $W = 1$  in these subplots. In the absence of buoyancy [see Fig. 1(b)], the  $W$  profiles change along the duct axis, beginning with a square shape at  $Z = 0$  and finally taking a parabolic shape further downstream. Therefore, maximum velocity always occurs at the centerline of  $Y = 0.5$ . Once buoyancy is initiated, the  $W$  profiles may shift towards the duct walls ( $Y = 0$  and  $1$ ), especially for the curves without radiation effect. Near the entrance (curve A), the velocity profile is fairly uniform over the cross-section. As the flow develops, the velocity in the core region is accelerated due to the entrance effect. Further downstream, the velocity profile becomes distorted with the maximum velocity towards the duct walls (curve C).

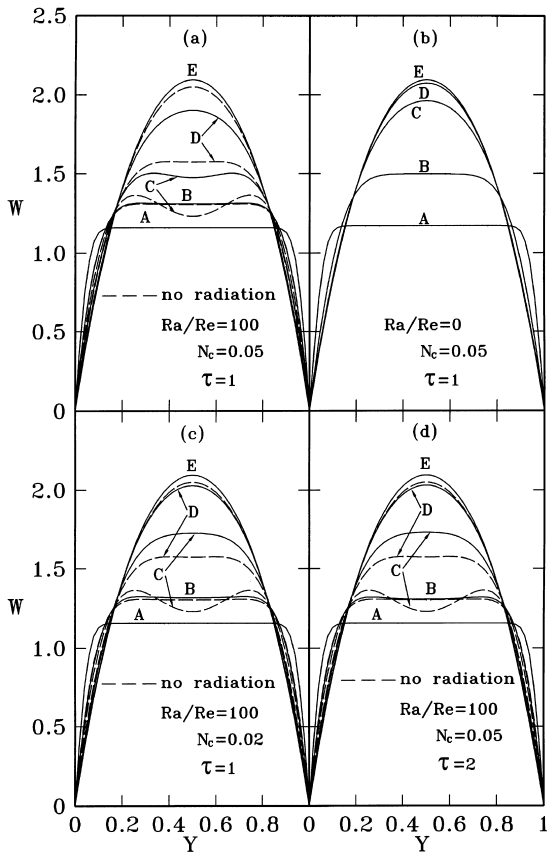


Fig. 1. Development of axial velocity profiles for different conditions: A:  $Z^* = 0.001$ , B:  $Z^* = 0.01$ , C:  $Z^* = 0.05$ , D:  $Z^* = 0.1$ , and E:  $Z^* = 0.3$ .

Finally, the velocity profiles become parabolic when the flow is fully developed. But as the radiation effect is taken into account (shown by the solid curves), the radiation tends to reduce the buoyancy effects. Therefore, the extent of the velocity shifting towards the walls becomes less significant, especially for stronger radiation cases (e.g., smaller  $N_c$  or larger  $\tau$ ). This can be understood by recognizing that the radiative heat flux is an additional mode of energy transport that causes a flatter temperature distribution. This would cause a small buoyancy effect. A close inspection of the solid and dashed curves discloses that the radiation effect tends to reduce  $W$  near the walls ( $Y = 0$  and  $Y = 1$ ). This implies that the results with radiation would reduce the friction factor.

Plotted in Fig. 2 are the temperature profiles at different axial locations, with the same legends as Fig. 1. In line with the wall heating, the temperature in the central portion of the duct increases gradually as the flow moves downstream. A comparison of the temperature profiles with and without radiation reveals that near the inlet the effect of radiation on the thermal development is insignificant.

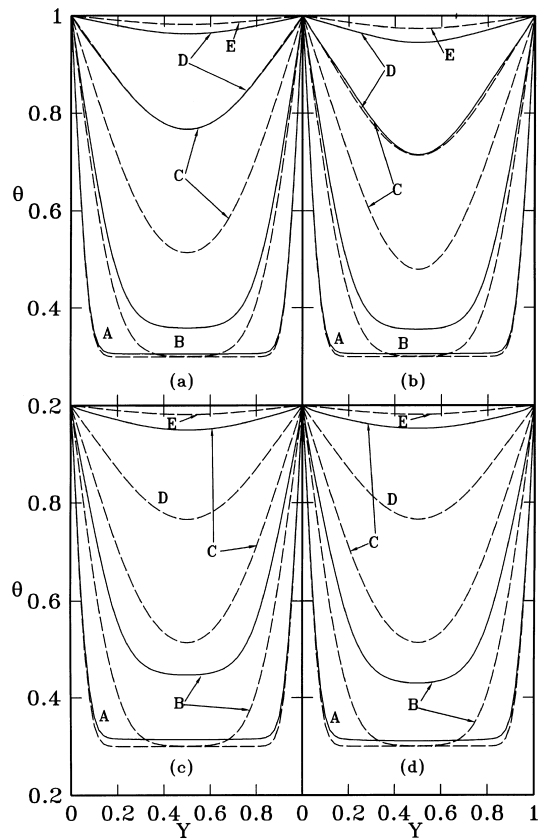


Fig. 2. Development of temperature profiles for different conditions: A:  $Z^* = 0.001$ , B:  $Z^* = 0.01$ , C:  $Z^* = 0.05$ , D:  $Z^* = 0.1$  and E:  $Z^* = 0.3$ .

But at the downstream location, radiation effect tends to equalize the temperature in the flow. A close inspection of Fig. 2(a) and (c) shows that as the condition-to-radiation parameter  $N_c$  decreases, the  $\theta$  becomes flatter. Additionally, the development of temperature profile is faster for a system with a stronger radiation effect ( $N_c = 0.02$ ). This can be understood by recognizing that the radiative heat flux is an additional mode of energy transport and as the magnitude of  $N_c$  decreases, the importance of the radiation term in Eq. (6) increases, relative to the conduction and convection. This causes an increase of the energy transport which leads to flatter temperature profiles. Similar result is also found for the case with a large  $\tau$ .

The effects of the ratio of Rayleigh number to Reynolds number  $Ra/Re$  and conduction-to-radiation parameter  $N_c$  on the axial variations of bulk temperature  $\theta_b$  are presented in Fig. 3. For comparison purposes, the results without thermal radiation effect are also plotted in Fig. 3, as shown by the dashed lines. It is clear that due to the stronger buoyancy effects, the development of bulk temperature is quicker for a sys-

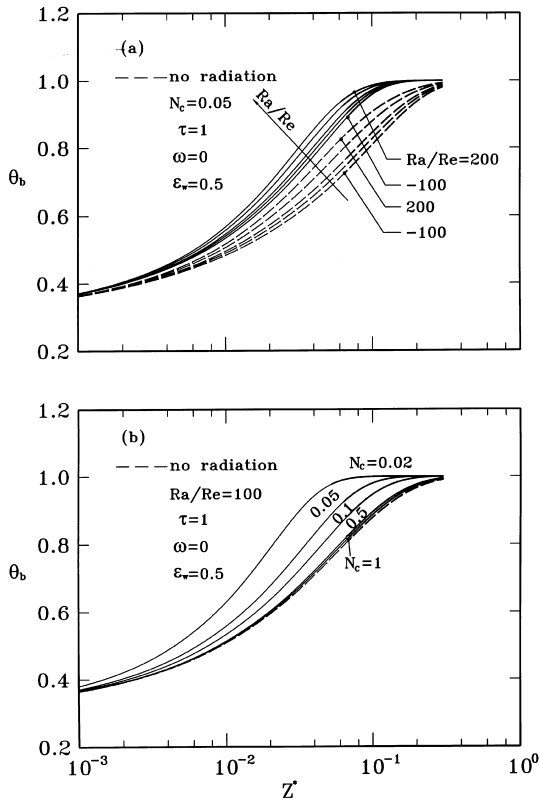


Fig. 3. Effects of  $Ra/Re$  and  $N_c$  on the axial variation of the bulk temperature  $\theta_b$  for  $\tau = 1$ ,  $\omega = 0$  and  $\epsilon_w = 0.5$ .

tem with a larger  $Ra/Re$ . In addition, the  $\theta_b$  is found to be essentially the same as that without radiation when the radiation-convection interaction is weak, such as for the curve of  $N_c = 1$ . However, when the interaction is intense ( $N_c = 0.02$ ), there is an essential difference. In contrast to the case without radiation, the radiation augments the rate of thermal development so that the  $\theta_b$  approaches the wall temperature ( $\theta = 1$ ) at a more rapid rate.

The effects of the conduction-to-radiation parameter  $N_c$  on the axial variations of local friction factor and Nusselt number are presented in Fig. 4. In Fig. 4, the total Nusselt number  $Nu_t$  with radiation is larger than that without radiation. This is due to two effects. The first is that radiation is an additional mechanism for heat transfer through the fluid, resulting in an increased heat flux. Secondly, the radiation source term augments the rate of thermal development so that the bulk temperature approaches the wall temperature at a more rapid rate. Both effects act to increase the local total Nusselt number  $Nu_t$ . Near the entrance, the  $Nu_t$  is raised only by the additional radiative heat flux, while further downstream, the more rapid thermal development contributes when the bulk temperature approaches the wall tem-

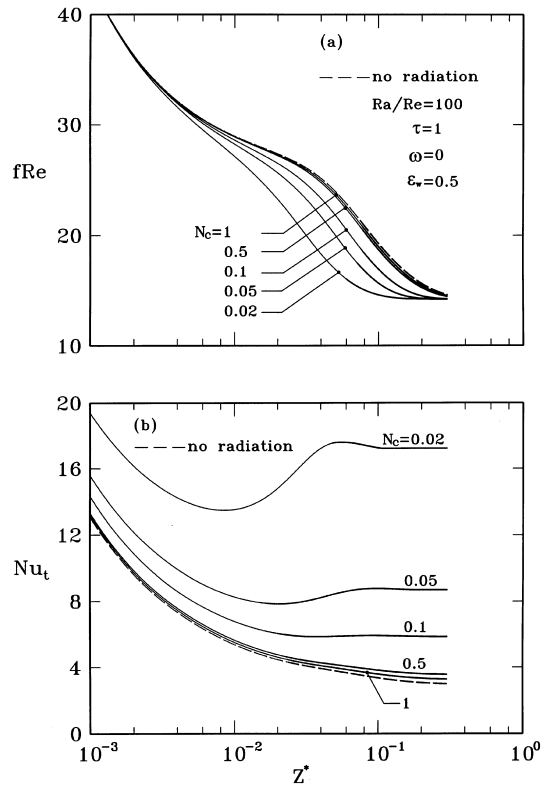


Fig. 4. Effects of  $N_c$  on the variation of the friction factor and the Nusselt number for  $Ra/Re = 100$ ,  $\tau = 1$ ,  $\omega = 0$  and  $\epsilon_w = 0.5$ .

perature. An overall inspection of Fig. 4 reveals that near the entrance, the local  $fRe$  fades away monotonically due to the entrance effect. And at these positions, the convective effect is predominant. Therefore, the effects of  $N_c$  on the local  $fRe$  are insignificant. But as the flow proceeds downstream, the radiation effect becomes important. The  $fRe$  decreases with a decrease in  $N_c$ . This is due to the fact that in the presence of radiation, the temperature field becomes flatter, as shown in Fig. 2. The buoyancy-assisting effect is thus reduced, which in turn, causes a reduction in  $fRe$  further downstream. But this is not the case with the local  $Nu_t$ . At downstream region, the  $Nu_t$  increases with a decrease in  $N_c$  due to the strong radiation effect.

The optical thickness  $\tau$  is another important factor which affects the heat transfer and fluid flow. Fig. 5 shows the effects of optical thickness  $\tau$  on the local  $fRe$  and  $Nu_t$ . When  $\tau$  is very small, the medium is almost transparent. Thus, it does not absorb and emit much energy. A large  $\tau$  corresponds to a strong radiatively participating medium. It is easy to see that there is more heat released from such a medium than from a weak radiatively participating medium. Thus, the local  $Nu_t$  increases with an increases in  $\tau$ . Additionally, the

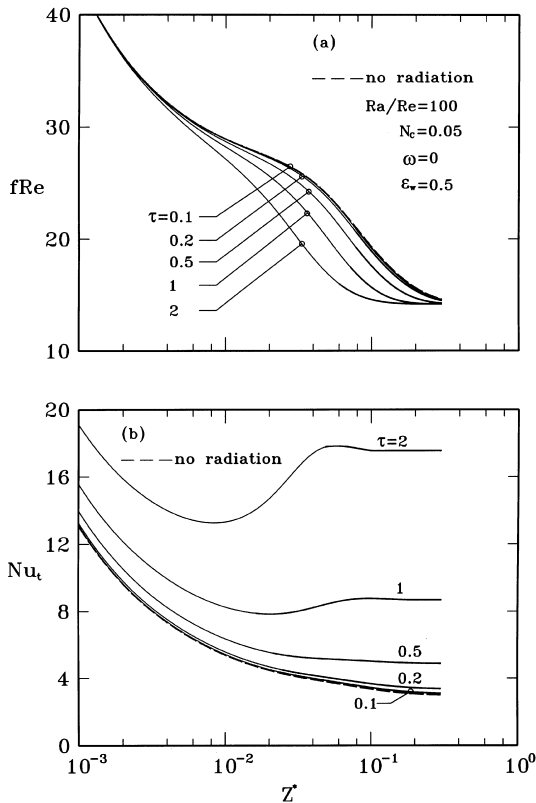


Fig. 5. Effects of  $\tau$  on the variation of the friction factor and the Nusselt number for  $Ra/Re = 100$ ,  $N_c = 0.05$ ,  $\omega = 0$  and  $\epsilon_w = 0.5$ .

radiation has a stronger effect on the local  $fRe$  and  $Nu_t$  for a system with a larger  $\tau$ , as compared with the results of no radiation.

In many applications, scattering processes are important in radiative heat transfer if particulates are present in the fluid. Therefore, the effects of the single scattering albedo on the fluid flow and heat transfer are of interest. For illustration, the scattering is assumed to be isotropic. Fig. 6 demonstrates the influence of scattering albedo  $\omega$  on the local  $fRe$  and  $Nu_t$  distributions. The  $Nu_t$  decreases with the increasing single scattering albedo  $\omega$ . This is expected, since as the scattering albedo  $\omega$  approaches zero, the emission and absorption of energy within the medium dominate. As the scattering albedo  $\omega$  approaches unity, scattering dominates and causes a small  $Nu_t$ . Like the results in Fig. 5(a), the local  $fRe$  is attenuated due to the radiation effect in the downstream region.

The effects of wall emissivity  $\epsilon_w$  on the local  $fRe$  and  $Nu_t$  are shown in Fig. 7. It is apparent that the local  $Nu_t$  without radiation is lower than that with radiation. Additionally, the local  $Nu_t$  increases with the increasing emissivity of the duct wall. Naturally, heat transfer is

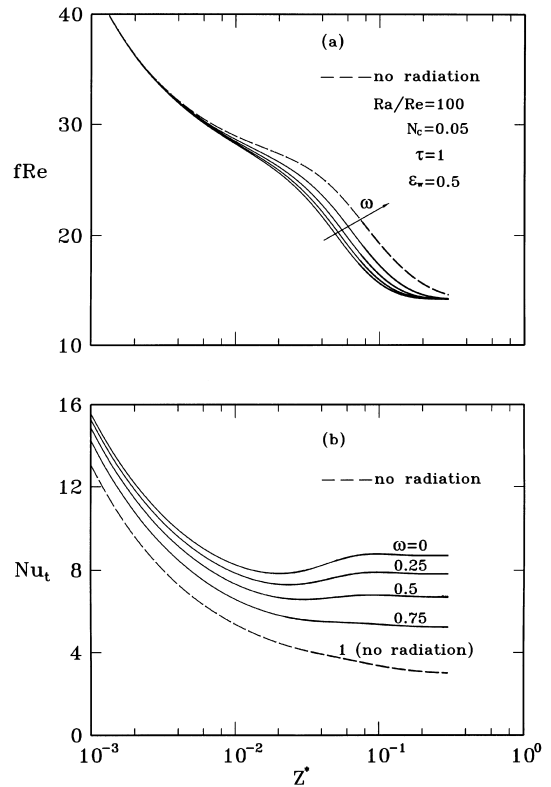


Fig. 6. Effects of  $\omega$  on the variation of the friction factor and the Nusselt number for  $Ra/Re = 100$ ,  $N_c = 0.05$ ,  $\tau = 1$  and  $\epsilon_w = 0.5$ .

maximum for a black duct (i.e.,  $\epsilon_w = 1$ ). It is also found that the stronger the radiation effect, the smaller the  $fRe$ .

In the conditions of high wall-to-coolant temperature differences, buoyancy may play a very critical role in the flow and heat transfer mechanism. Hence, it is interesting to examine its influence on the flow and heat transfer. Fig. 8 shows the effects of the  $Ra/Re$  on the local  $fRe$  and  $Nu_t$  for  $N_c = 0.05$ ,  $\tau = 1$ ,  $\omega = 0$  and  $\epsilon_w = 0.5$ . Positive  $Ra/Re$  indicates a buoyancy-assisting flow, while negative  $Ra/Re$  means a buoyancy-opposing flow. It is obvious that the buoyancy-assisting effect increases the  $fRe$  and  $Nu_t$  as compared with the results of buoyancy-free case ( $Ra/Re = 0$ ), and the extent of increase in  $Nu_t$  increases with the magnitude of  $Ra/Re$ . But for the case of negative  $Ra/Re$ , the opposing buoyancy decreases the local  $fRe$  and  $Nu_t$ . Additionally, relative to the results without radiation, the radiation effect would decrease the local  $fRe$  for the buoyancy-assisting case. But for buoyancy-opposing flow, the presence of radiation would enhance the local  $fRe$ . These are due to the fact that the radiation tends to reduce the buoyancy effects.



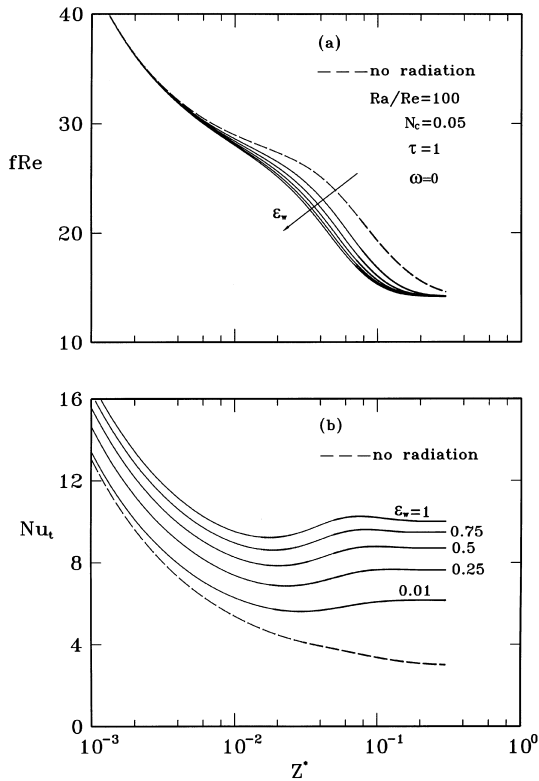


Fig. 7. Effects of  $\epsilon_w$  on the variation of the friction factor and the Nusselt number for  $Ra/Re = 100$ ,  $N_c = 0.05$ ,  $\tau = 1$  and  $\omega = 0$ .

**5. Conclusions**

In the present work, the problem of mixed convection flow and heat transfer in a vertical square duct with consideration of thermal radiation has been studied numerically. A novel vorticity–velocity method successively solved the three-dimensional parabolic governing equations. The radiative transfer equation was solved by the discrete ordinates method. The effects of the Rayleigh number-to-Reynolds number parameter  $Ra/Re$ , conduction-to-radiation parameter  $N_c$ , optical thickness  $\tau$ , single scattering albedo  $\omega$  and wall emissivity  $\epsilon_w$  on the flow and heat transfer are examined in detail. What follows is a brief summary.

1. In the presence of radiation, the thermal development develops at a more rapid rate relative to that without radiation.
2. The local friction factor  $fRe$  and Nusselt number  $Nu_t$  are enhanced with an increase in the magnitude of  $Ra/Re$  for buoyancy-assisting case.
3. The local total Nusselt number  $Nu_t$  is augmented by the radiation effects. The  $fRe$  is reduced (increased) for the buoyancy-assisting (opposing) flow as the thermal radiation effect is taken into account.

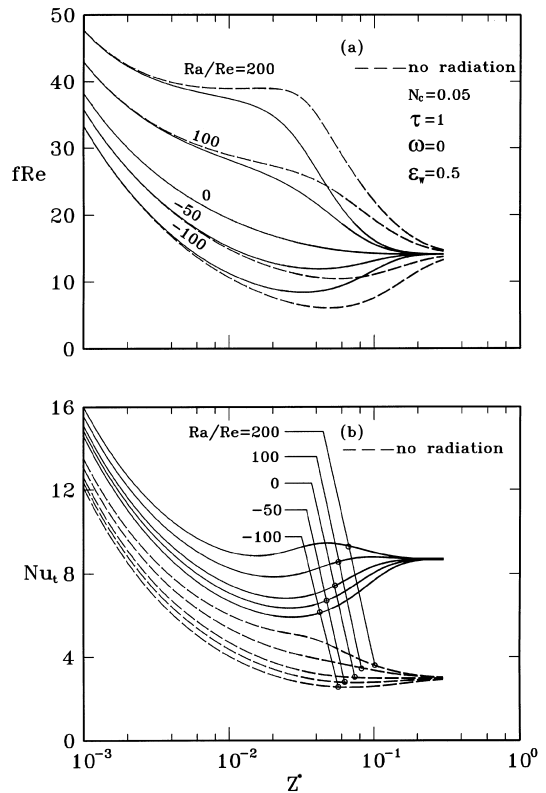


Fig. 8. Effects of  $Ra/Re$  on the variation of the friction factor and the Nusselt number for  $N_c = 0.05$ ,  $\tau = 1$ ,  $\omega = 0$  and  $\epsilon_w = 0.5$ .

4. In contrast to the results without radiation, radiation effects augment the  $Nu_t$ . And the extent of enhancement in  $Nu_t$  increases with a decrease (increase) in  $N_c$  or  $\omega$  ( $\tau$  or  $\epsilon_w$ ).
5. The laminar case is treated in this work and intended to provide a first step towards future work which will investigate the radiation effects on turbulent mixed convection in vertical ducts.

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